

Lower Hybrid Drift Waves in the Magnetotail

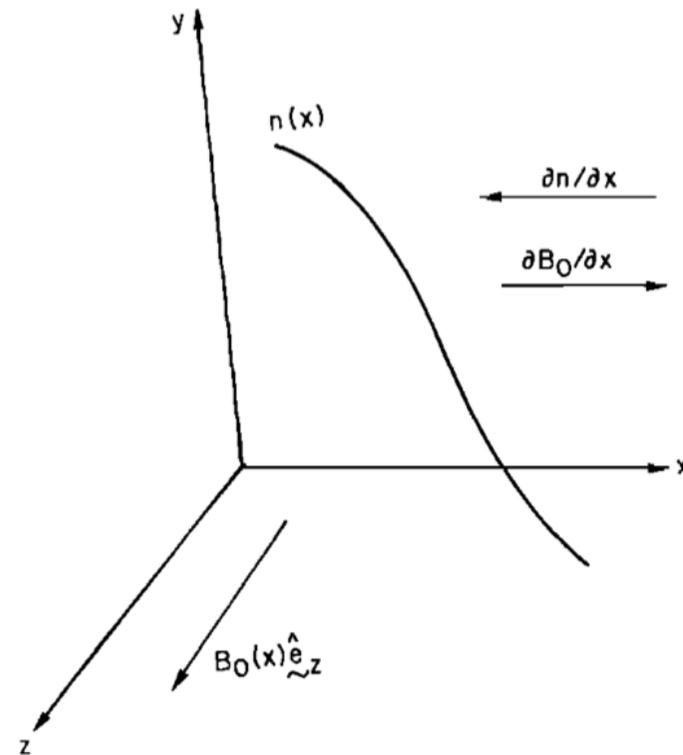
Master thesis by Cecilia Norgren
Supervised by Andris Vaivads, Yuri and Mats

I will talk about...

- Lower hybrid drift waves
- ... in the magnetosphere
- The Cluster data
- Observations
- Summary

Lower Hybrid Drift Waves

- Energy stored in plasma inhomogeneities



Lower Hybrid Drift Waves

- Energy stored in plasma inhomogeneities
- Ions drift due to density gradients and couple with waves

$$v_{Di} = T_i / eBL_n$$

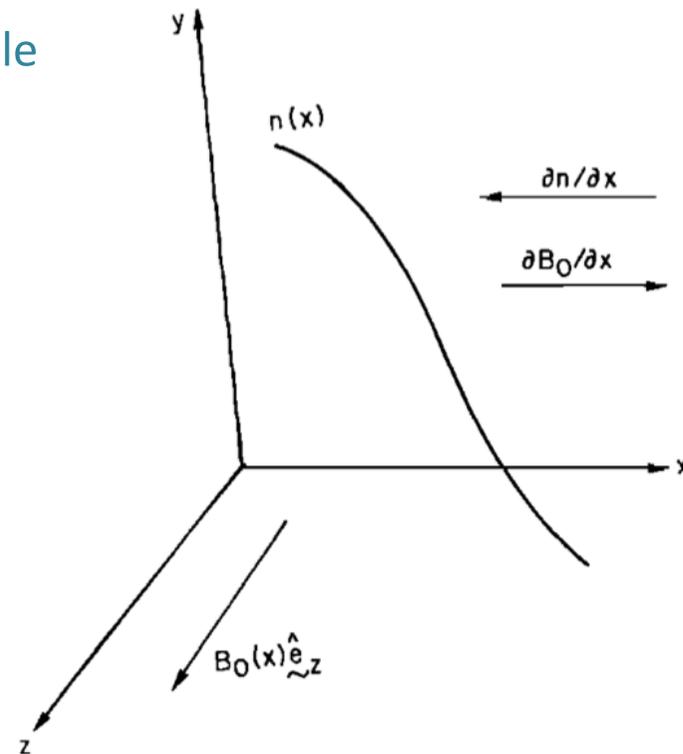
- Ions are unmagnetized
- Electrons are magnetized

$$\omega_{ci} \ll \omega \ll \omega_{ce}$$

$$\rho_e \ll D < \rho_i$$

- Perpendicular waves

$$k_{\perp} \gg k_{\parallel}$$



Lower Hybrid Drift Waves

- Excitation condition

$$\frac{L_n}{\rho_i} < \left(\frac{m_i}{m_e}\right)^{\frac{1}{4}} \approx 7 \quad L_n = \left(\frac{1}{n} \frac{\partial n}{\partial x}\right)^{-1}$$

- Wavelength

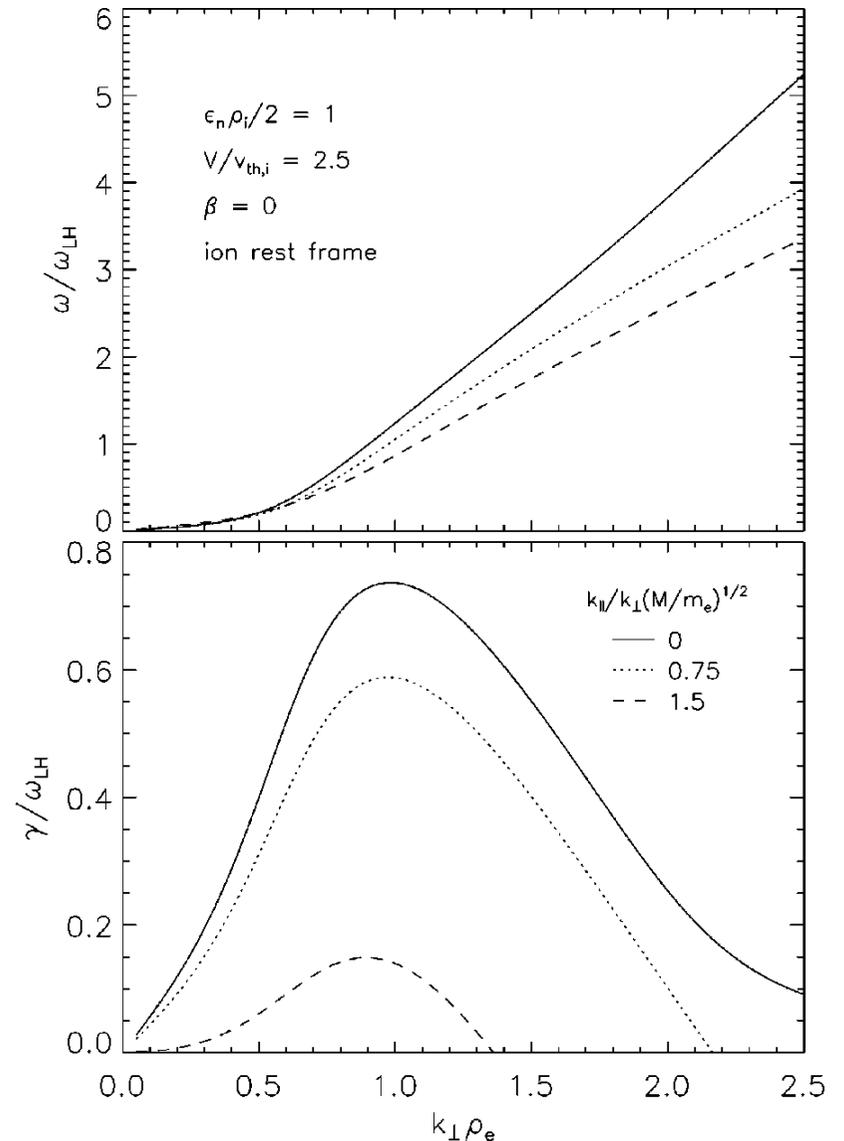
$$k_{\perp} \rho_e \approx 1 \quad \Rightarrow \quad \lambda_{LH} = 2\pi / k_{\perp} \approx 2\pi \rho_e$$

- Frequency

$$\omega = \omega_{LH} = \sqrt{\Omega_e \Omega_i}$$

- Large growth rate

➤ Nonlinear wave can change rapidly!



Lower Hybrid Drift Waves

- In low β -plasma they are generally electrostatic
 - Finite plasma β has a stabilizing effect on the electric wave mode
 - As β increases, a longer wavelength electromagnetic mode becomes prominent
- Role in magnetic reconnection?

Why do we want to measure them?

- We want to know what they look like
- Might provide insight into general electron scale dynamics
- Might play an important role in various plasma processes

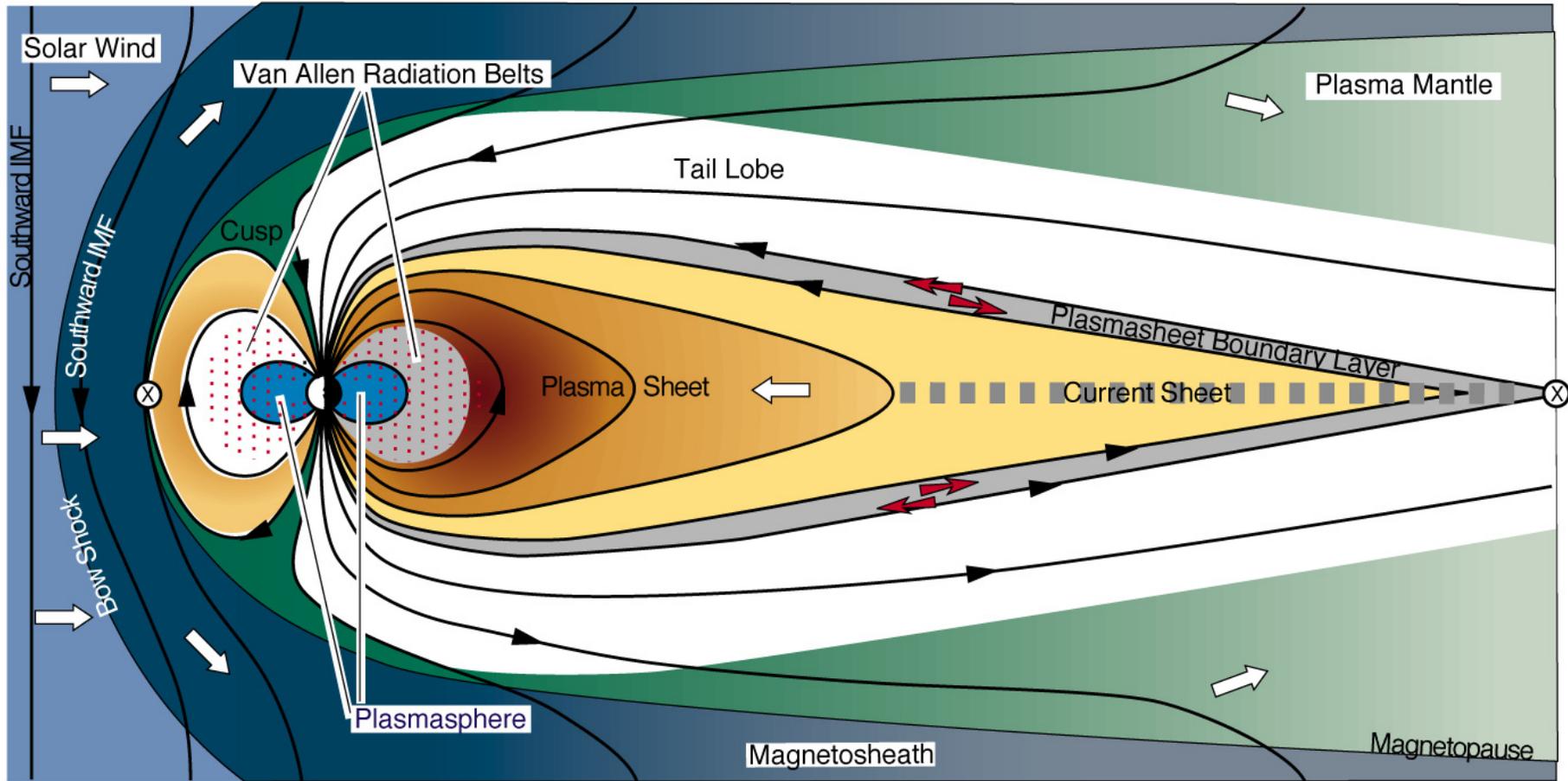
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What do we want to measure?

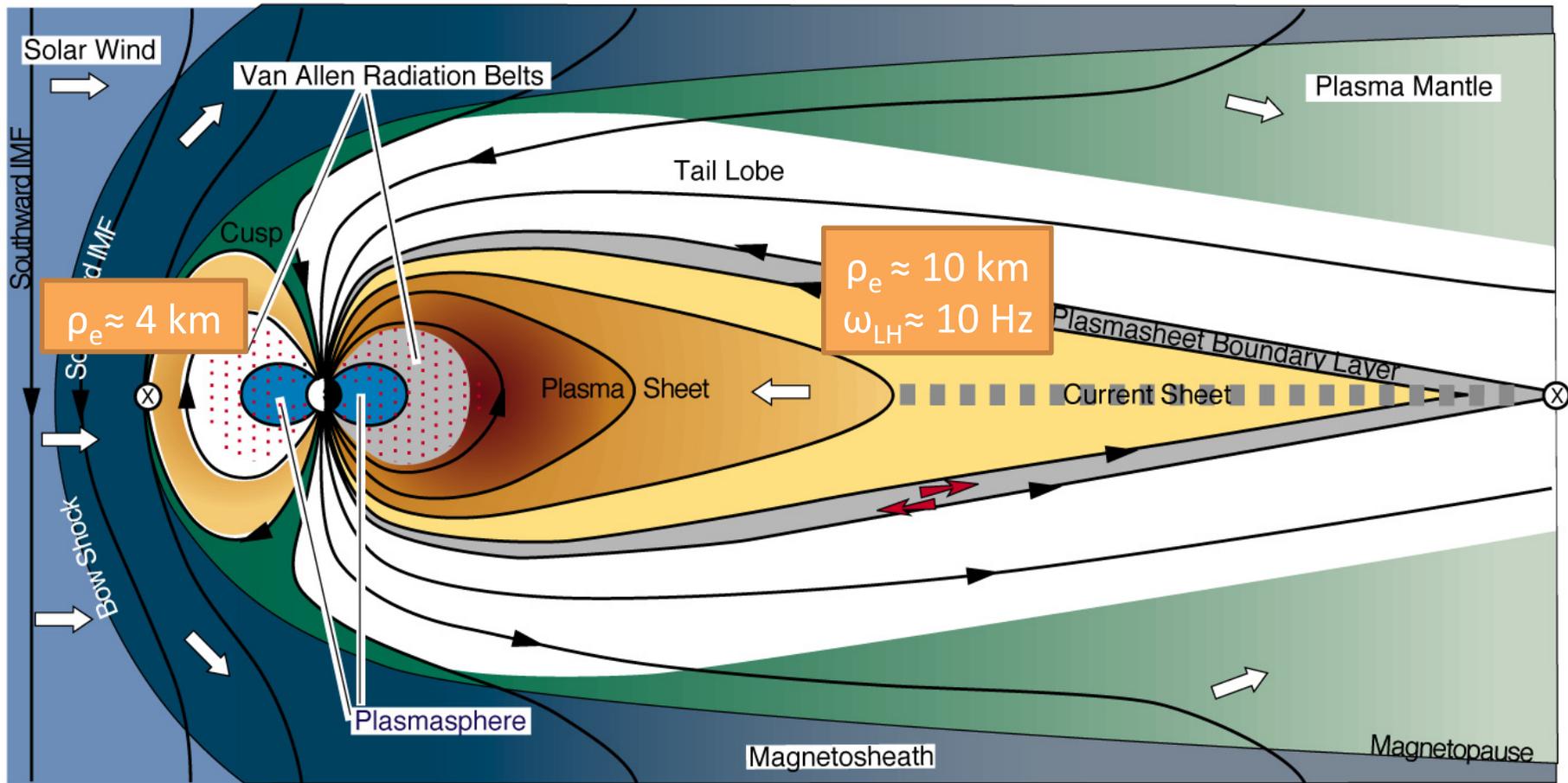
- Velocity
- Wavelength
- Potential
- Instability conditions

The Magnetosphere



The Magnetosphere

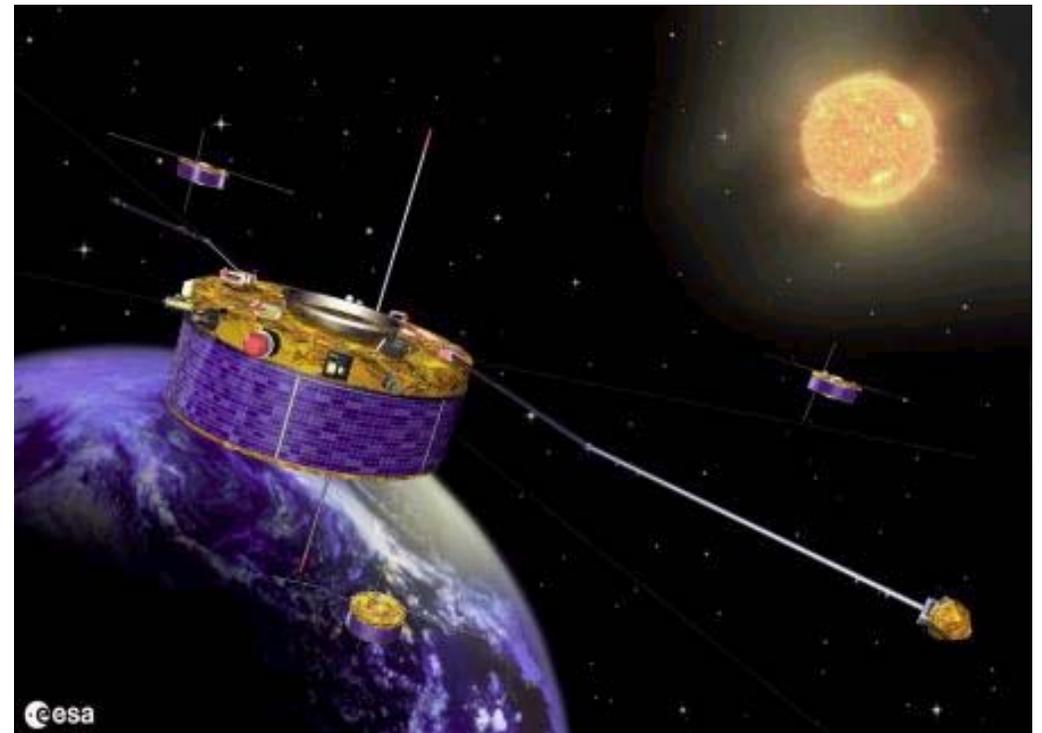
- Short separation in the magnetotail + burst mode!



Cluster

2007

- C3, C4 separated by 40 km in the magnetotail
- 14 planned boundary layer burst modes
 - ✓ 3 real plasma sheet boundary layer crossings



Space vs. laboratory

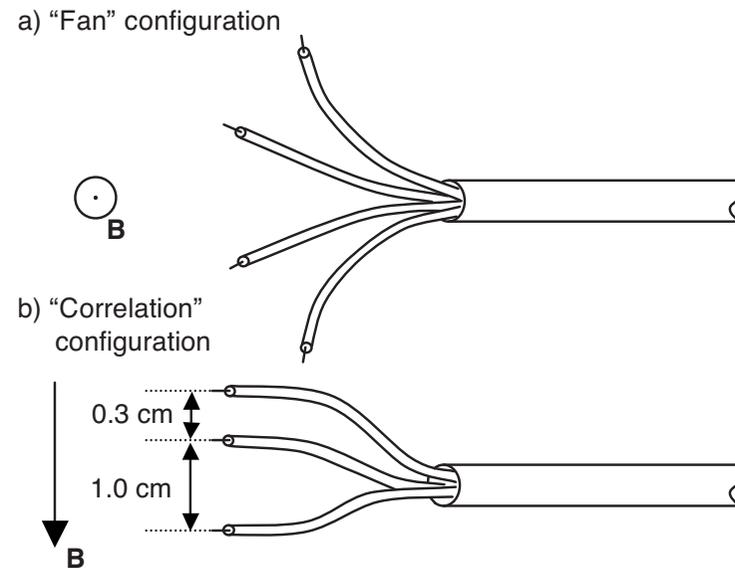
Wouldn't it be easier to just set up a laboratory experiment?

Cluster is so small in the big magnetosphere:

$\lambda_{LH} = 60 \text{ km}$
88 m between probes
separation 10 km

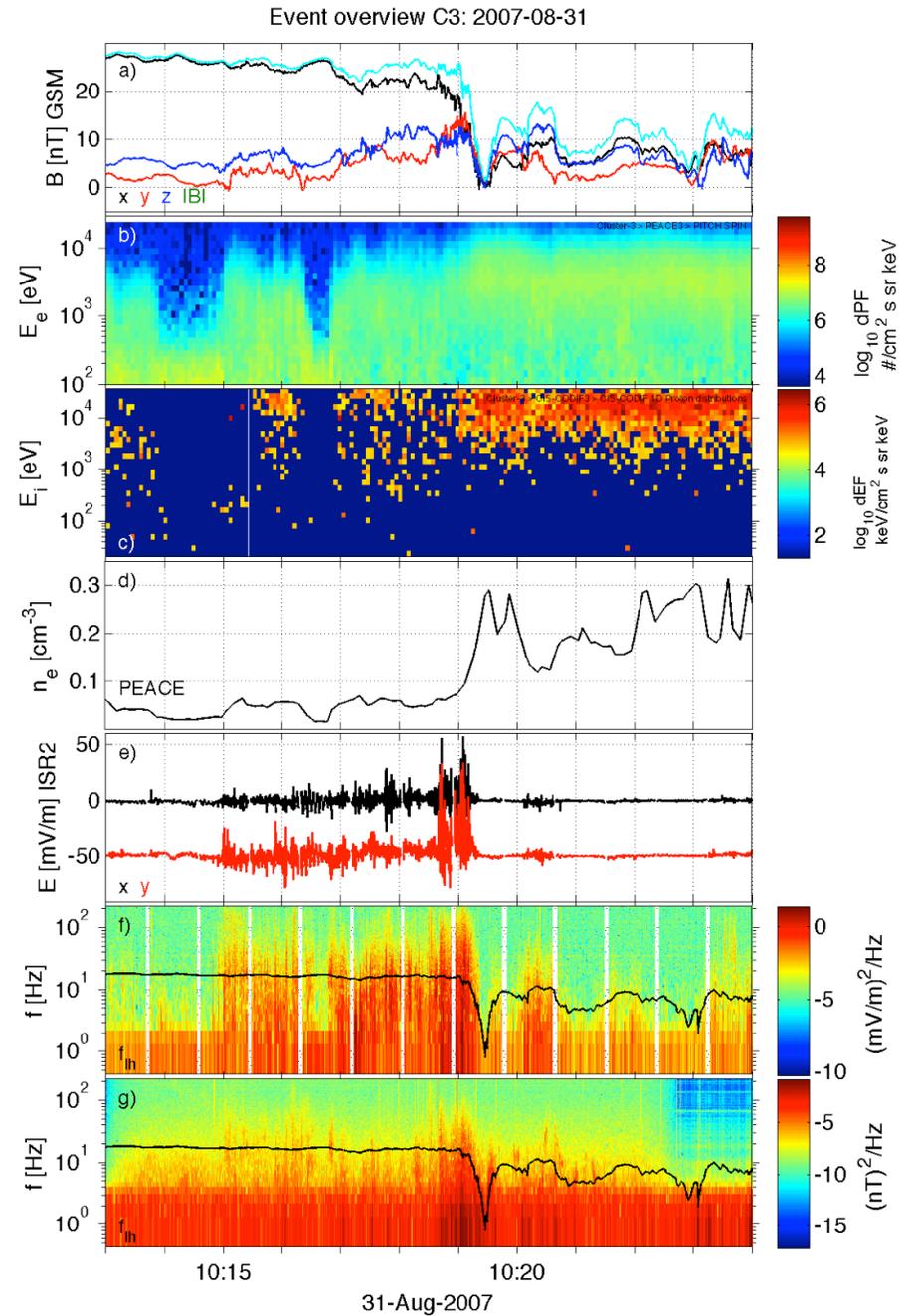
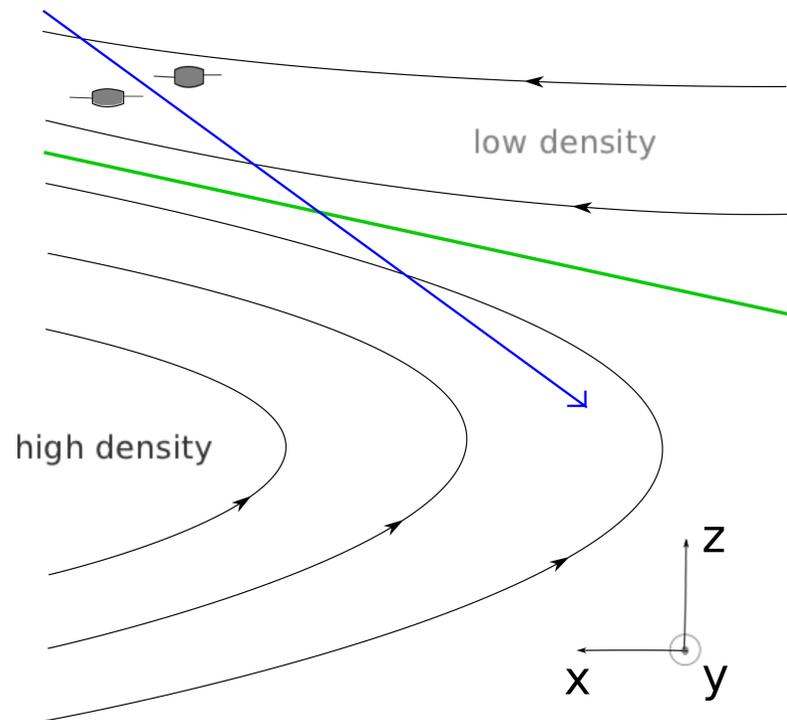
In laboratory (Fox 2010):

$\lambda_{LH} = 1.2 \text{ mm}$
probe width 0.3 mm
separation 3 mm



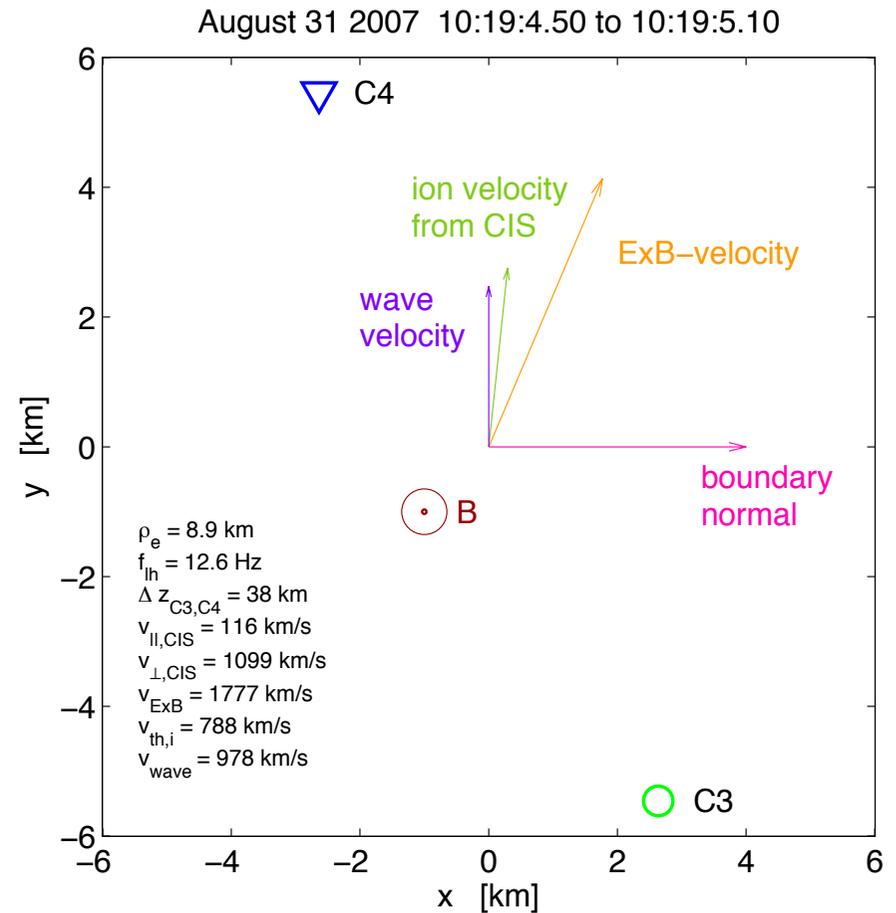
August 31, 2007

Tail lobe → Plasma sheet



The event set up

- B along z
 - Minimum variance analysis gives the normal direction b (x)
 - Only direction left is the wave propagation direction! (y)
- $\Delta y \approx 11$ km
- $\lambda_{LH} \approx 2\pi\rho_e \approx 56$ km

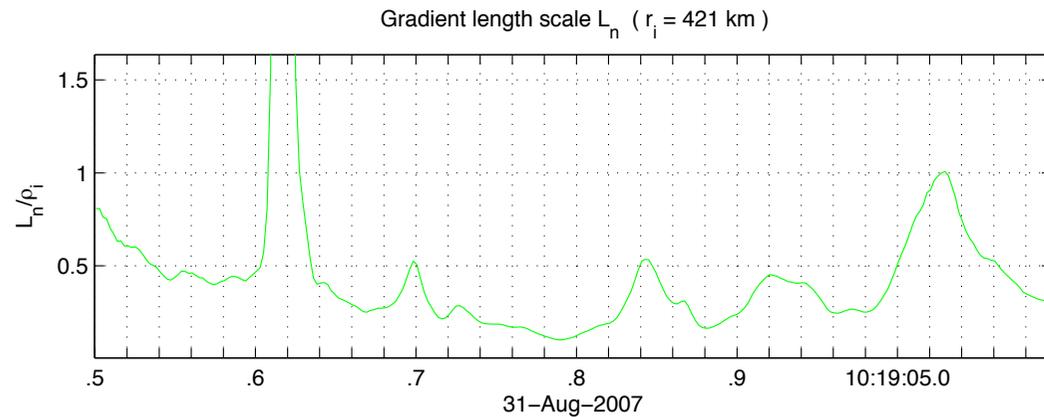


Are the conditions right?

We have a density gradient, but is it sharp enough?

$$\frac{L_n}{\rho_i} < \left(\frac{m_i}{m_e} \right)^{\frac{1}{4}} \approx 7 \quad L_n = \left(\frac{1}{n} \frac{\partial n}{\partial x} \right)^{-1}$$

- Assumed total pressure balance

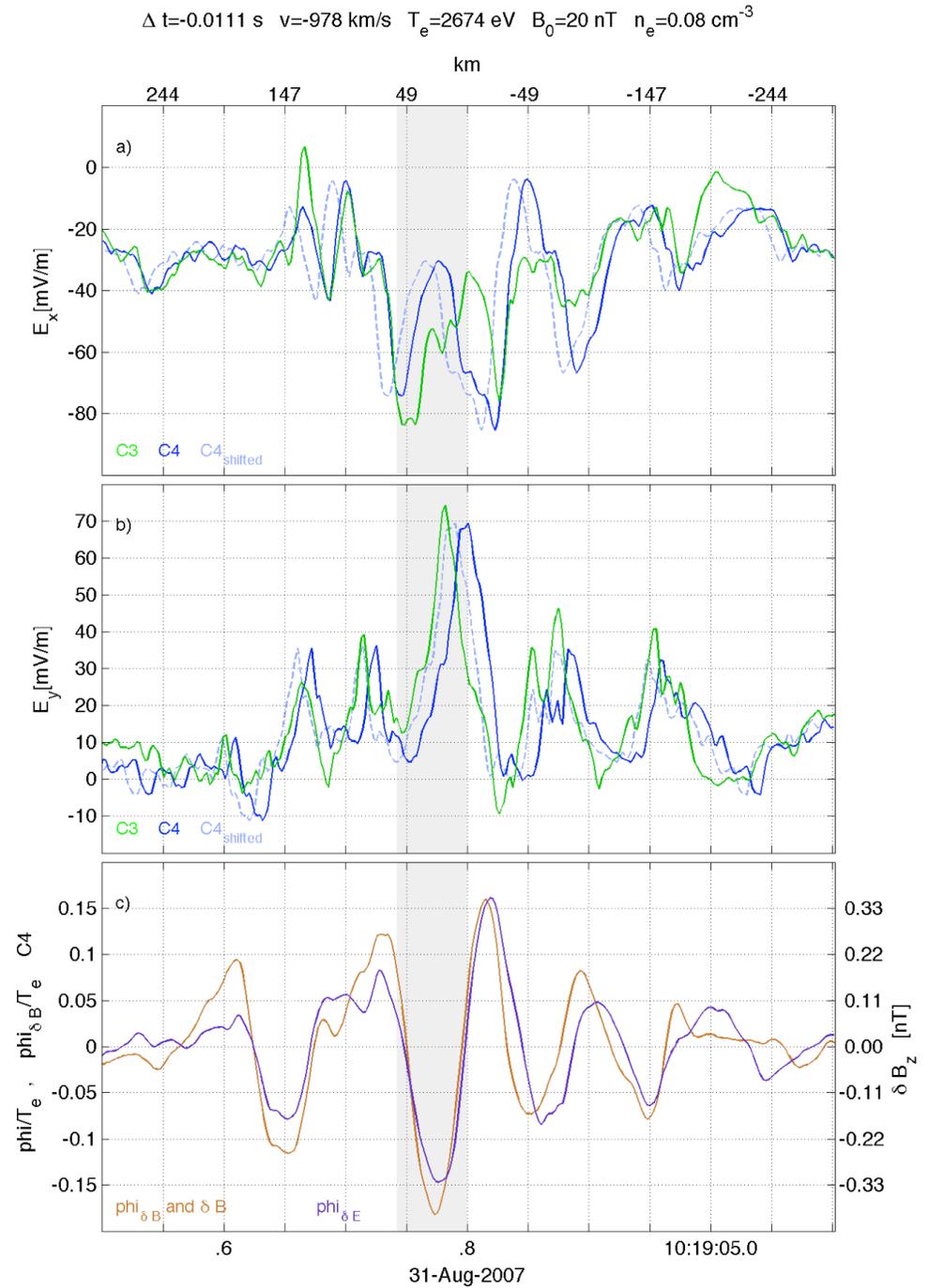


- Integration of ExB normal velocity from DC E-field gives $L_n / \rho_i \approx 5$

δE and ϕ

- We look how much the electric field is delayed between C3 and C4

- $v = 978 \text{ km/s}$
- $\lambda \approx 80 \text{ km}$



δE and ϕ

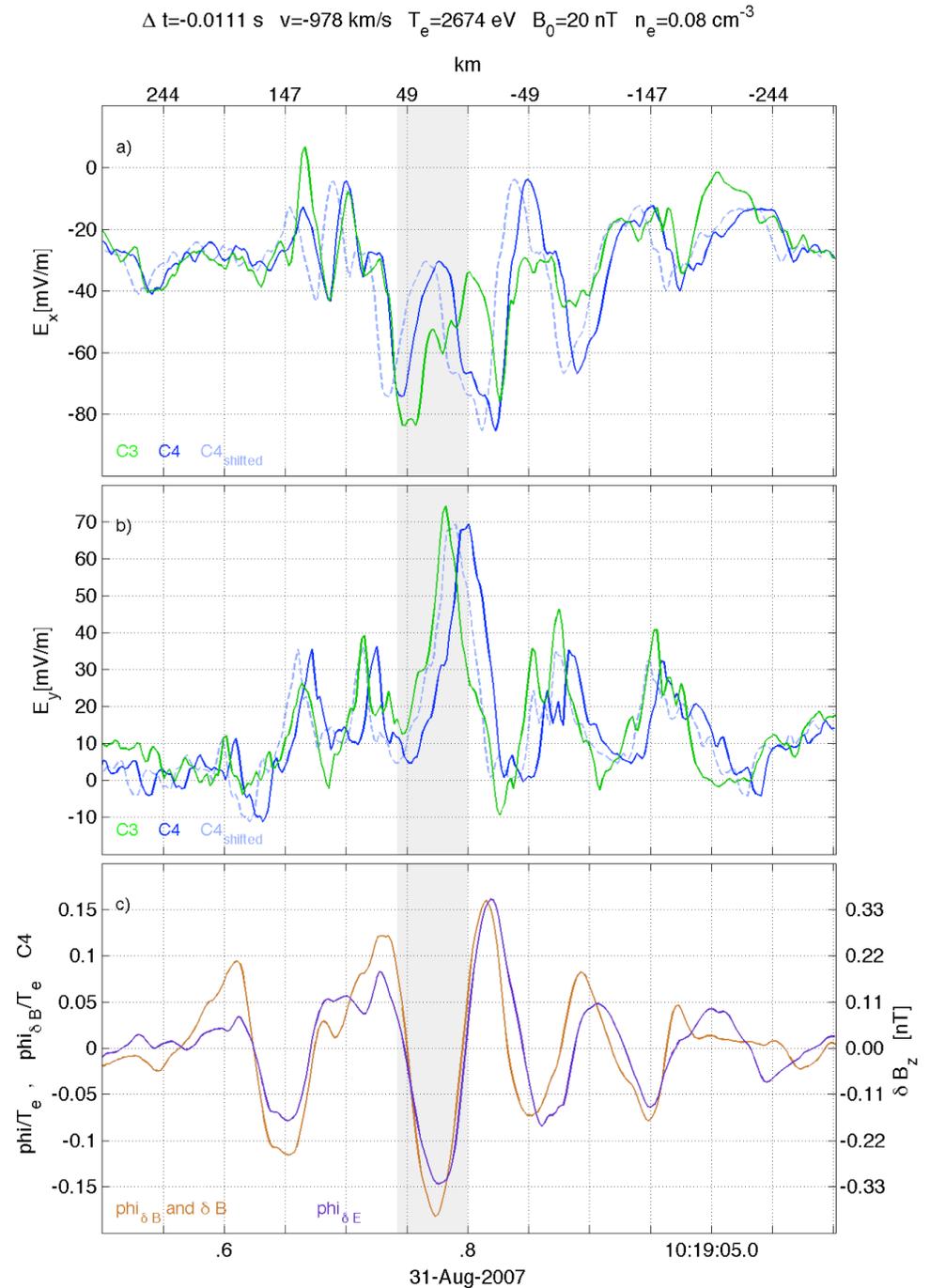
- We look how much the electric field is delayed between C3 and C4

- $v = 978 \text{ km/s}$
- $\lambda \approx 80 \text{ km}$

- The electrostatic potential:

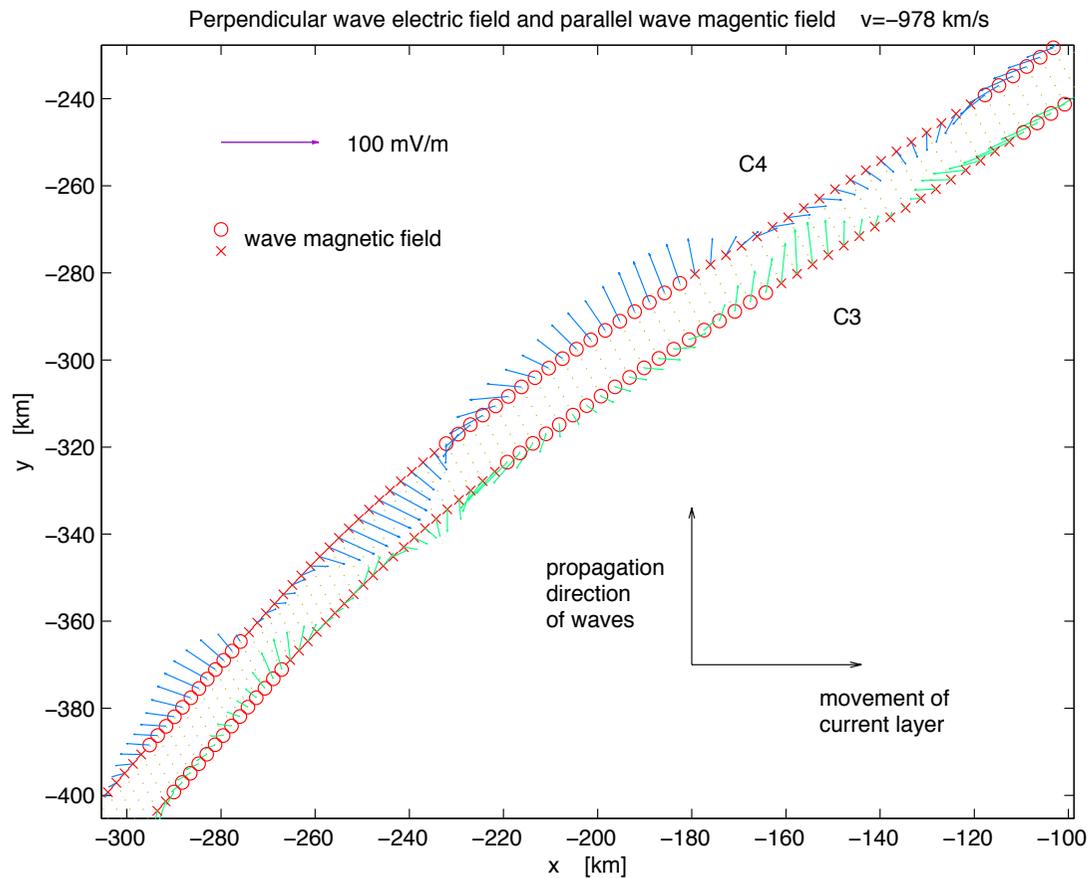
$$\phi = \int \vec{E} dt \cdot \vec{v} \quad \text{Integrate!}$$

- $e\phi$ is 10-30% of $k_B T_e$
The potential might affect the electrons!



δE in field aligned coordinate system

- ✓ δE_{perp} for each time step
- ✓ δB_{\parallel} for each time step
 - Repetitive pattern!



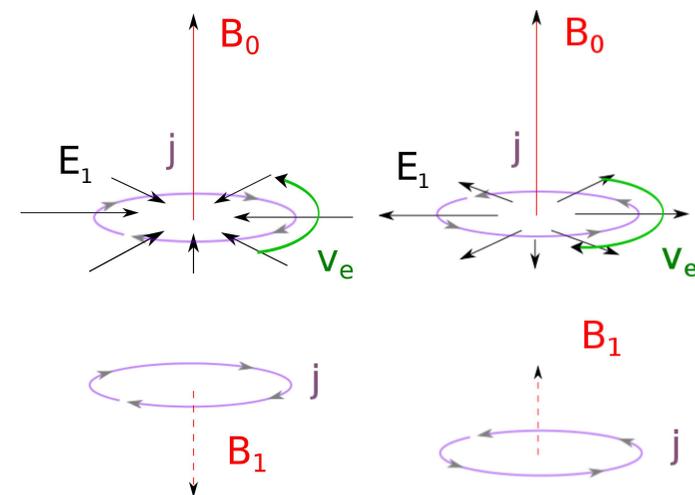
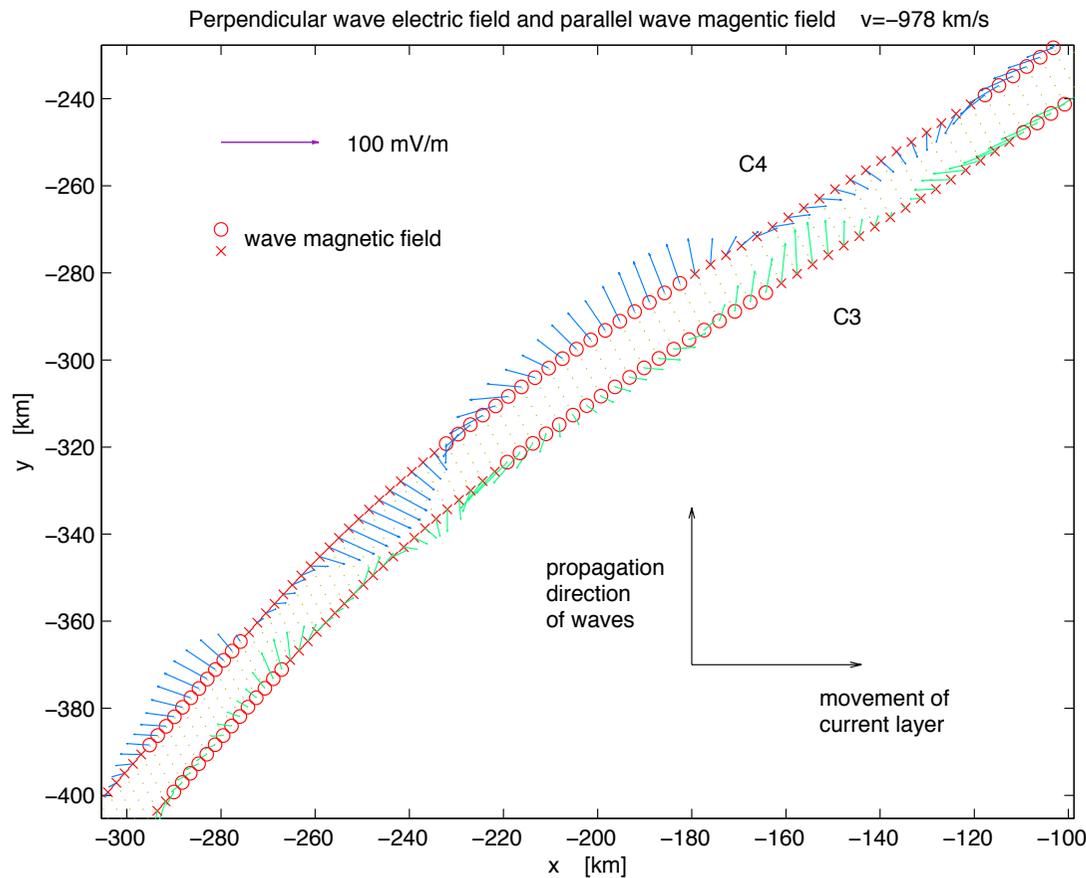
δE in field aligned coordinate system

- ✓ δE_{perp} for each time step
- ✓ δB_{\parallel} for each time step
 - Repetitive pattern!

Ions unmagnetized:

$$j = ne(v_i - v_e) = -ne \frac{E_1 \times B}{B^2}$$

$$\nabla \times B_1 = \mu_0 j$$

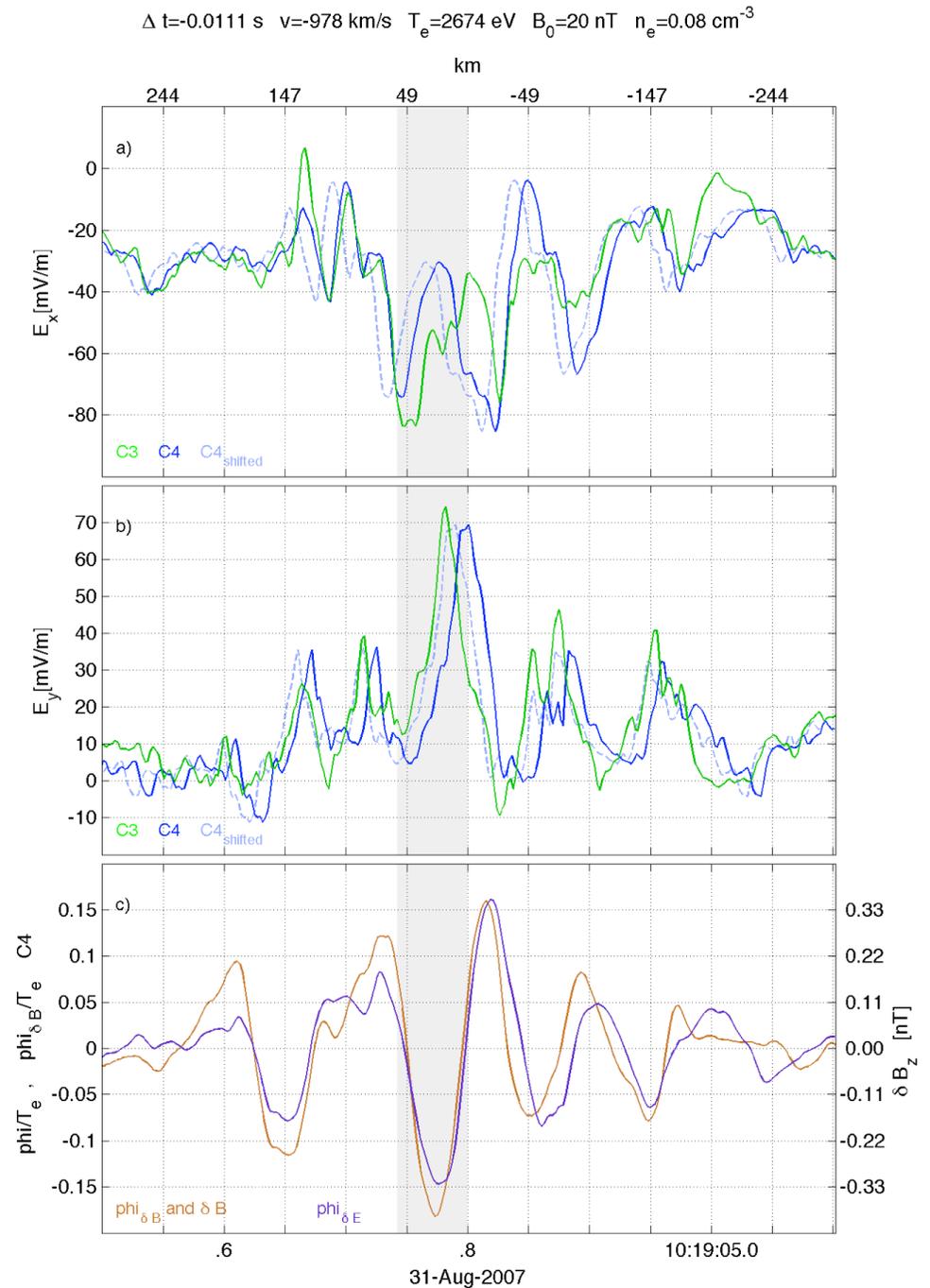


δB and ϕ

- The wave magnetic field is linearly related to the electrostatic potential

$$\phi_{\delta B} = \frac{B}{\mu_0 n e} \delta B$$

(via the reasoning on the last slide...)



δB and ϕ

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$$\phi_{\delta B} = \frac{B}{\mu_0 n e} \delta B$$

(via the reasoning on the last slide...)

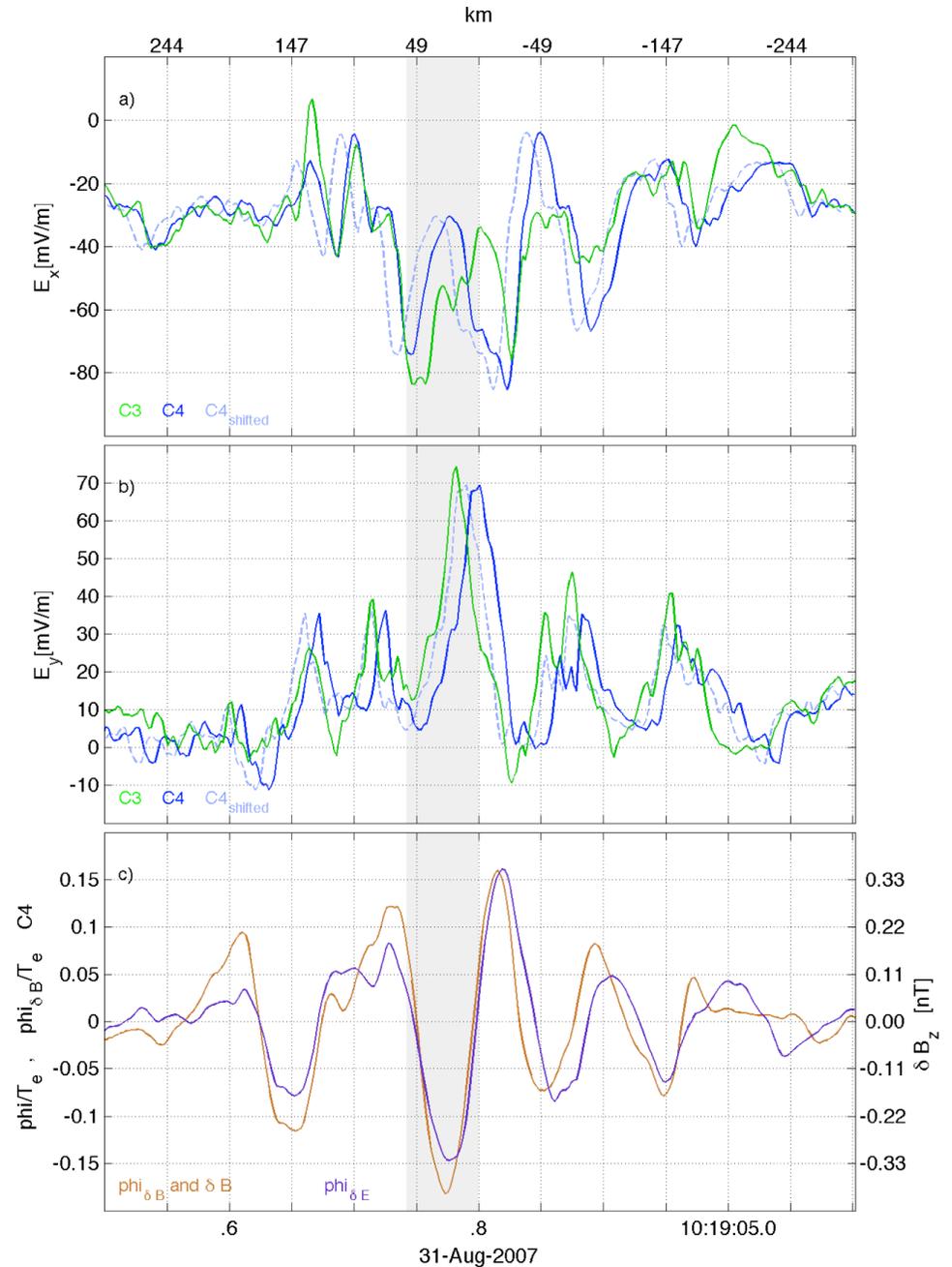
✧ Very good correspondance!

✓ Verification

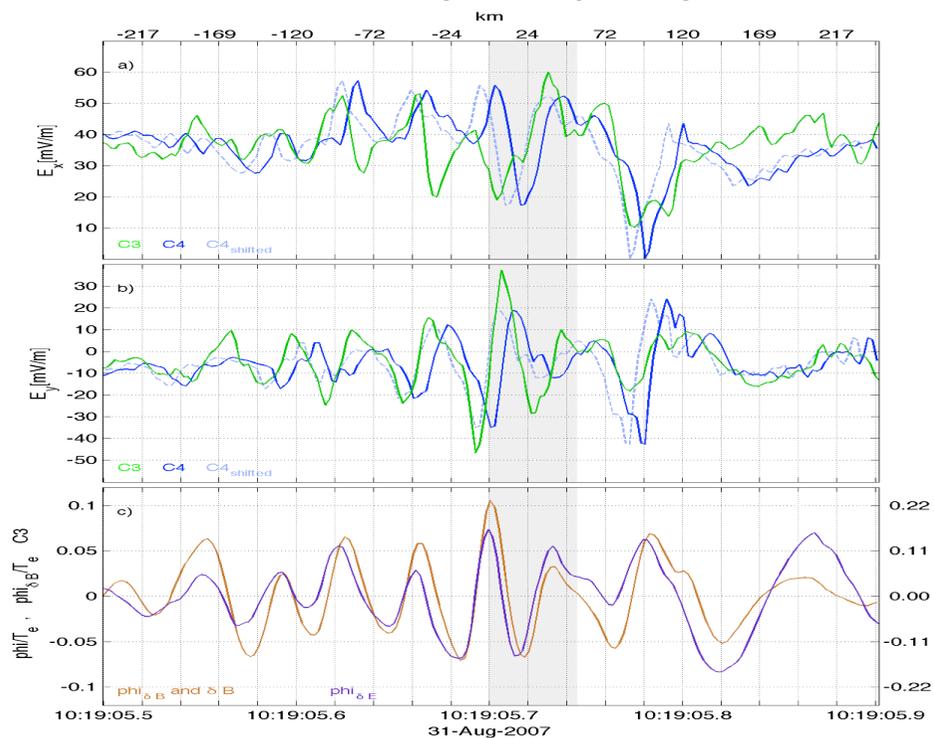
✓ Tool

Now we can estimate ϕ even in cases when cross spacecraft correlation is not possible.

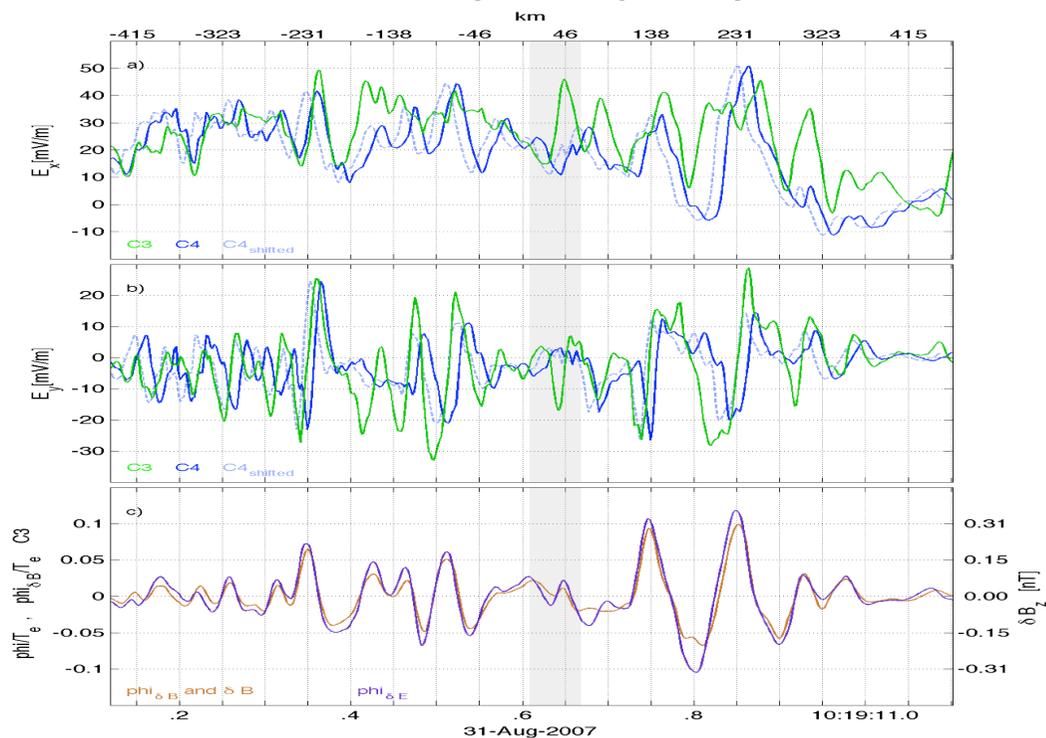
$\Delta t = -0.0111$ s $v = -978$ km/s $T_e = 2674$ eV $B_0 = 20$ nT $n_e = 0.08$ cm⁻³



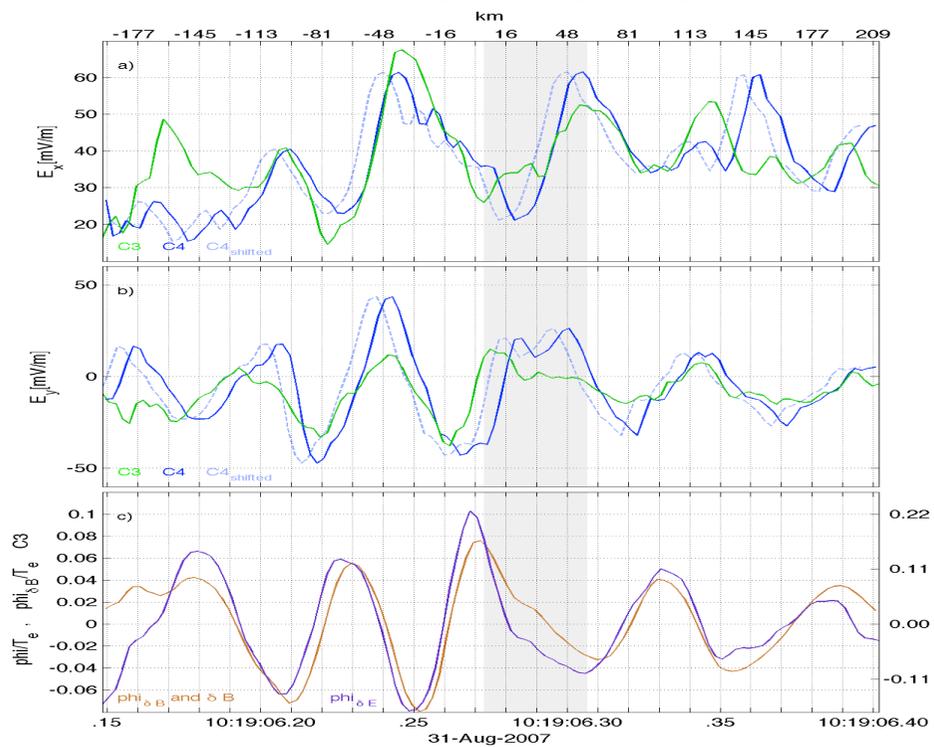
$\Delta t = -0.0079$ s $v = 1204$ km/s $T_e = 2708$ eV $B_0 = 20$ nT $n_e = 0.08$ cm $^{-3}$



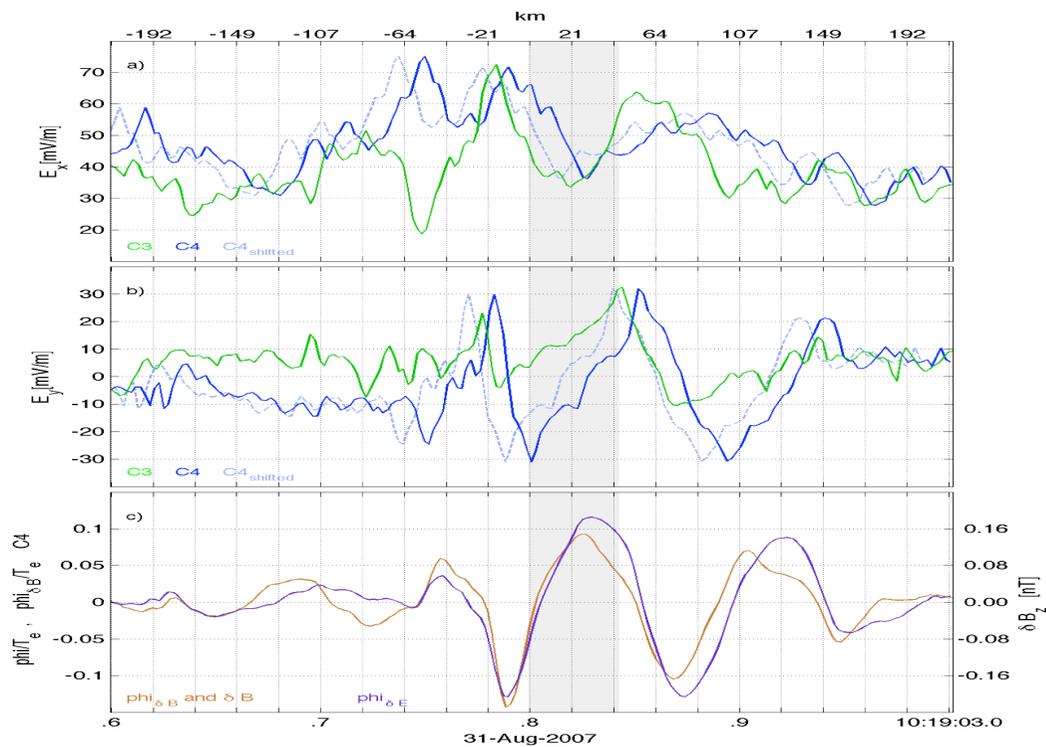
$\Delta t = -0.0126$ s $v = 922$ km/s $T_e = 2860$ eV $B_0 = 21$ nT $n_e = 0.11$ cm $^{-3}$



$\Delta t = -0.0051$ s $v = 1610$ km/s $T_e = 2730$ eV $B_0 = 20$ nT $n_e = 0.08$ cm $^{-3}$



$\Delta t = -0.0123$ s $v = 1067$ km/s $T_e = 2609$ eV $B_0 = 24$ nT $n_e = 0.07$ cm $^{-3}$



δB and δE

Electrostatic wave or not?

Faraday's law

$$\frac{|E_1|}{|B_1|} = \frac{\omega}{k} \quad \text{electromagnetic wave}$$

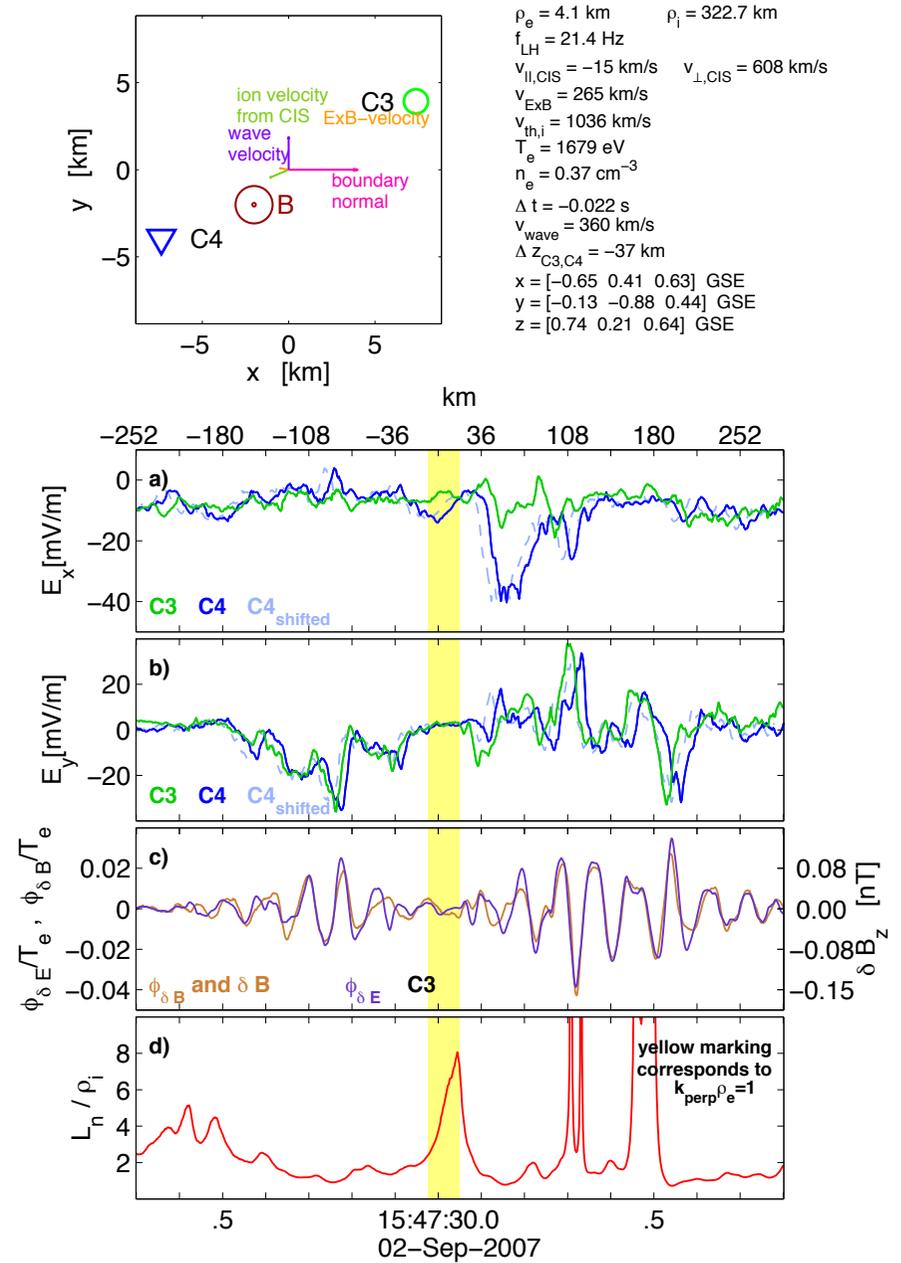
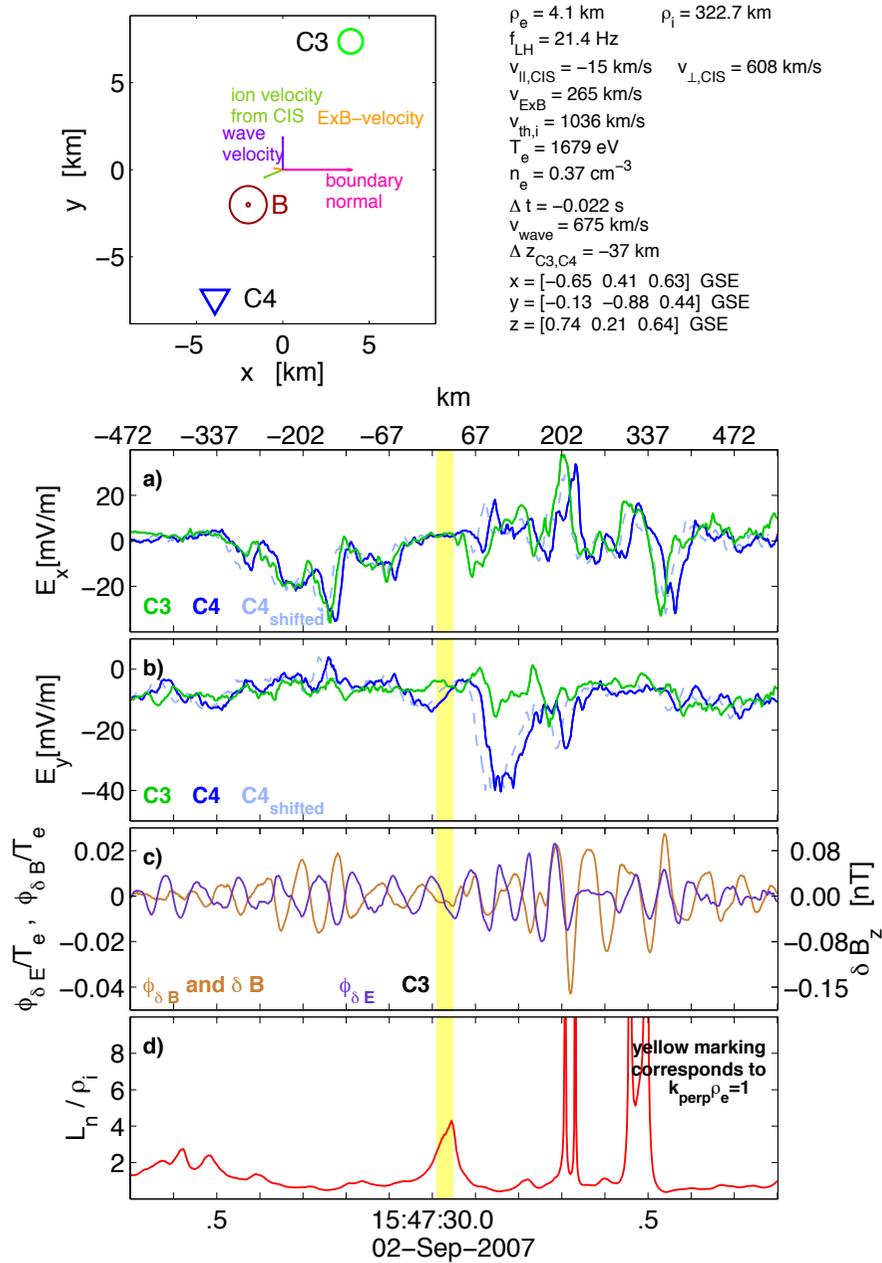
$$\frac{|E_1|}{|B_1|} \gg \frac{\omega}{k} \quad \text{electrostatic wave}$$

We have

$$\frac{|E_1|}{|B_1|} = \frac{60 \text{ mV} / \text{m}}{0.6 \text{ nT}} = 10^5 \text{ km} / \text{s}$$

MVA n

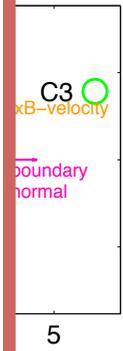
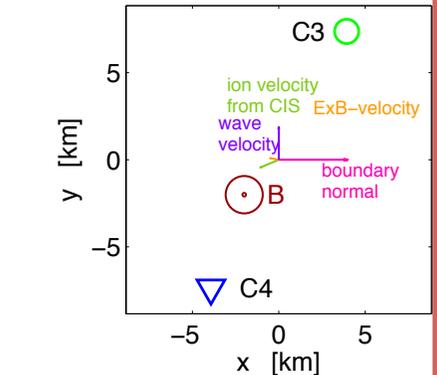
cheating n



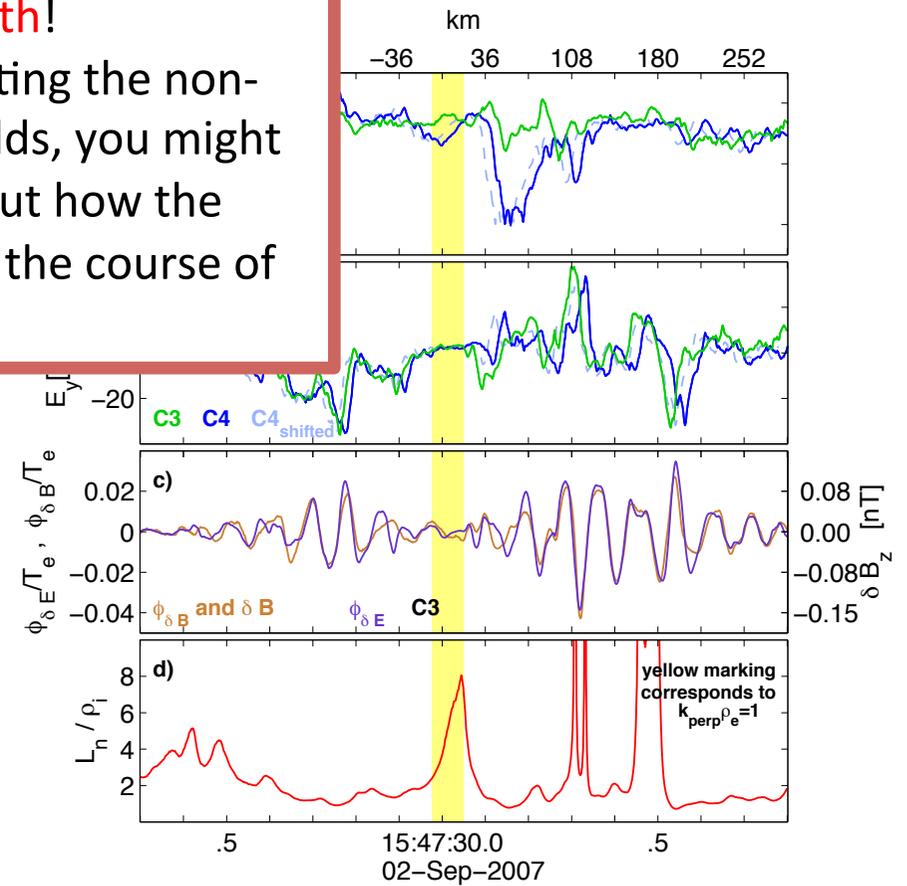
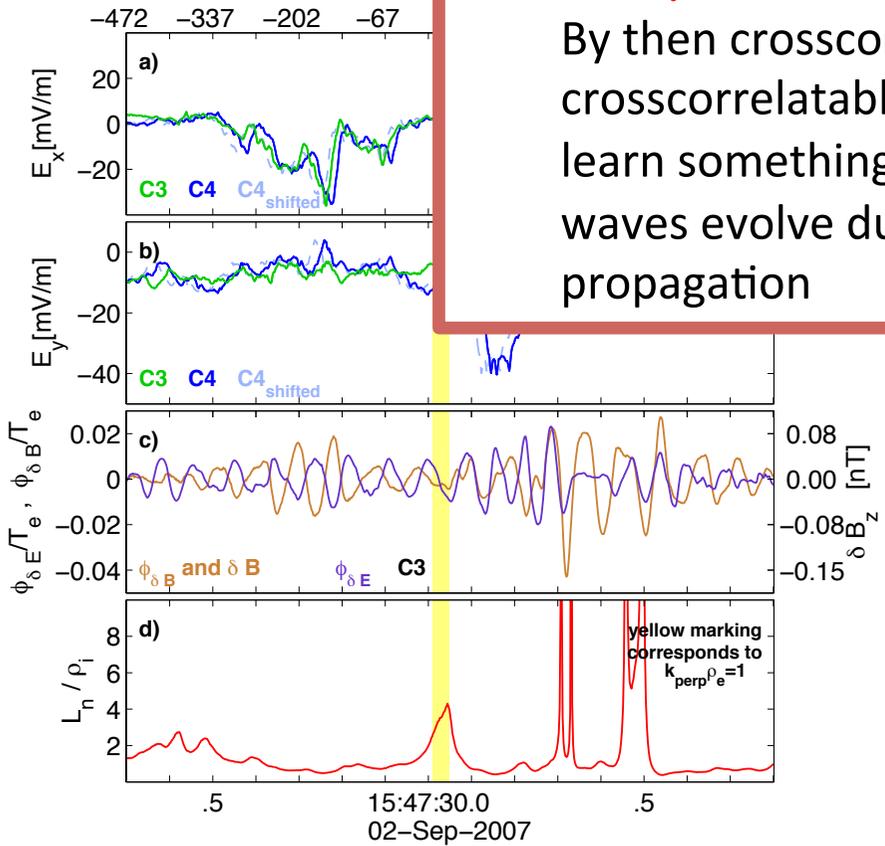
MVA n

cheating n

- Define z along B
 - Try different **propagation directions** until you obtain the good waveform
 - Try different **velocities** until you get the correct amplitude
- **Velocity and wavelength!**
- By then crosscorrelating the non-crosscorrelatable fields, you might learn something about how the waves evolve during the course of propagation



$\rho_e = 4.1 \text{ km}$ $\rho_i = 322.7 \text{ km}$
 $f_{LH} = 21.4 \text{ Hz}$
 $V_{||,CIS} = -15 \text{ km/s}$ $V_{\perp,CIS} = 608 \text{ km/s}$
 $V_{ExB} = 265 \text{ km/s}$
 $v_{th,i} = 1036 \text{ km/s}$
 $T_e = 1679 \text{ eV}$
 $n_e = 0.37 \text{ cm}^{-3}$
 $\Delta t = -0.022 \text{ s}$
 $V_{wave} = 360 \text{ km/s}$
 $\Delta z_{C3,C4} = -37 \text{ km}$
 $x = [-0.65 \ 0.41 \ 0.63] \text{ GSE}$
 $y = [-0.13 \ -0.88 \ 0.44] \text{ GSE}$
 $z = [0.74 \ 0.21 \ 0.64] \text{ GSE}$



What have we not done?

- Saturation mechanisms: Why are the waves not more turbulent?
 - Current relaxation
 - Plateau formation
- Wavelength dependence
- ϕ in a wider parameter space using δB

Summary

Short separation between spacecraft + burst mode + boundary layer

➤ Three events in all Cluster data!

- Velocity: $v_{\text{wave}} \approx 1000 \text{ km/s}$
 - Wavelength: $\lambda_{\text{measured}} \approx 50\text{-}100 \text{ km}$
 - Gradient length scale: $L_n/\rho_i \approx 0.5$
 - Electrostatic potential: ϕ is 10-40 % of T_e
 - Wave magnetic field: $\phi_{\delta B} \approx \phi_{\delta E}$
 - Structure of wave: Vortices!
- $v_{\text{wave}} \approx v_{\text{ions}}$
 $\lambda_{\text{measured}} \approx \lambda_{\text{theory}}$
 $\%v_{Di}/v_i \approx \rho_i/2L_n \approx 1$